直线加速 Kinnersley 黑洞中 Dirac 粒子的热辐射*

(重庆三峡学院物理与电子工程学院,重庆 404000) (2007年4月19日收到 2007年6月14日收到修改稿)

在直线加速 Kinnersley 时空中,将相互耦合的 Dirac 方程化为二阶方程,采用新的乌龟坐标变换,在视界面附近 消除二阶方程中的耦合化成了标准波动方程,得到辐射温度函数和 Hawking 热辐射谱.

关键词:黑洞, Dirac 方程, 乌龟坐标变换, Hawking 辐射 PACC: 9760L, 0420

1.引 言

自从 Hawking 指出黑洞会发射热辐射以来 ,人 们在黑洞热力学的研究上做了大量工作,特别是赵 峥等人^[1-3]在 Damour-Ruffini 工作的基础上发展出 来的广义乌龟坐标变换法,在研究各类动态黑洞的 Hawking 辐射上获得了很大的成功.但该方法用于研 究动态轴对称黑洞和任意加速 Kinnersley 黑洞中 Dirac 粒子的 Hawking 辐射时遇到了困难,原因是在 这两类黑洞中相互耦合的 Dirac 方程不能经过分离 变量法而完全退耦,则在视界面附近不能化成单一 分量的标准波动方程.10年以后,吴双清和蔡勖等 人[4-10]建议了处理该问题的一种方法,他们对一阶 和二阶 Dirac 方程同时作乌龟坐标变换处理 然后利 用一阶方程提供的关系式去消除二阶方程的相互耦 合项,使得每一个分量满足的二阶方程在视界面附 近都能化为单一分量的标准波动方程,成功地处理 了这两类黑洞中旋量粒子的热辐射问题 这种方法 简洁明了 具有普适性.文献 11 对电磁直线加速动 态黑洞中 Dirac 粒子的 Hawking 辐射作了探讨,得到 了相应的结果,但其对 Dirac 方程处理的方法较为繁 杂 不具有一般性,而文献 12 采用的是不同的乌龟 坐标变换对同一课题作了探讨,讨论不是完整的,部 分结论也不成立.

本文用吴双清和蔡勖建议的方法和新的乌龟坐 标变换^[13],讨论直线加速 Kinnerslev 黑洞中荷电 Dirac 粒子的 Hawking 辐射,在视界面附近将相互耦合的 Dirac 方程进行了完全退耦,化成为单一分量的标准波动方程,得到黑洞辐射温度函数,导出Hawking 热谱公式.

2. 直线加速 Kinnersley 的时空线元和 零标架的选取

作直线加速 Kinnersley 的时空线元为

$$ds^{2} = g_{00}dv^{2} + 2g_{01}dvdr + 2g_{02}dvd\theta$$

$$+ g_{22}d\theta^{2} + g_{33}d\varphi^{2}, \qquad (1)$$

式中

$$g_{00} = 1 - \frac{2M}{r} - 2 a r \cos \theta + \frac{Q^2}{r^2}$$

- $4 a \frac{Q^2}{r} \cos \theta - r^2 f^2$,
$$g_{01} = g_{10} = -1$$
, (2)
$$g_{02} = g_{20} = -r^2 f$$
,
$$g_{22} = -r^2$$
,
$$g_{33} = -r^2 \sin^2 \theta$$
.

而 $f = -a \sin \theta$ 参量 M = M(v)和 Q = Q(v)分别是 黑洞的质量和所带电荷 a = a(v)是加速度参量.

容易计算出度规行列式和非零逆变分量为

$$\begin{split} g &= -r^* \sin^2 \theta , \\ g^{01} &= g^{10} = -1 , \\ g^{11} &= -\left[1 - \frac{2M}{r} - 2ar \cos \theta + \frac{Q^2}{r^2} - 4a \frac{Q^2}{r} \cos \theta\right] , \end{split}$$

杨波

^{*} 重庆市教育委员会科学技术项目计划(批准号:KJ071111)资助的课题.

$$g^{12} = g^{21} = f ,$$

$$g^{22} = -\frac{1}{r^2} ,$$
(3)

 $g^{33} = -\frac{1}{r^2 \sin^2 \theta}.$

令黑洞的事件视界面方程为

$$H = H(v, r, \theta) = 0$$

或

$$r_{H} = r_{H} (v, \theta), \qquad (4)$$

则(4)式应该满足零曲面方程条件:

$$g^{\mu\nu} \frac{\partial H}{\partial x^{\mu}} \frac{\partial H}{\partial x^{\nu}} = 0.$$
 (5)

由(4)式和(5)式可得

$$\frac{\partial H}{\partial r}\frac{\partial r}{\partial v} - \frac{\partial H}{\partial v} = 0,$$

$$\frac{\partial H}{\partial r}\frac{\partial r}{\partial \theta} - \frac{\partial H}{\partial \theta} = 0,$$
(6)

由(3)(5)和(6)式可得事件视界 r_H = r_H(v,θ)方程 满足

$$1 - \frac{2M}{r_{H}} - 2ar_{H}\cos\theta + \frac{Q^{2}}{r_{H}^{2}}$$
$$- 4a\frac{Q^{2}}{r_{H}}\cos\theta - 2r_{Hv} + 2fr_{H\theta} + \frac{r_{H\theta}^{2}}{r_{H}^{2}} = 0 , \quad (7)$$

 $\mathbf{H}\mathbf{P} \mathbf{r}_{Hv} = \partial r/\partial v |_{r=r_H} \mathbf{r}_{H\theta} = \partial r/\partial \theta |_{r=r_H}.$

由度规分量(2)式选取如下的零标架协变分量: $l_{\mu} = \begin{bmatrix} 1 & 0 & 0 & 0 \end{bmatrix}$, $n = \begin{bmatrix} r & /2 & -1 & -r^2 f & 0 \end{bmatrix}$

$$m_{\mu} = \begin{bmatrix} g_{00}/2 & -1 & -r \end{bmatrix} = 0],$$

$$m_{\mu} = \frac{r}{\sqrt{2}} \begin{bmatrix} 0 & 0 & 1 & i \sin\theta \end{bmatrix},$$
 (8)

$$\overline{m}_{\mu} = \frac{r}{\sqrt{2}} \begin{bmatrix} 0 & 0 & 1 & -i \sin\theta \end{bmatrix},$$

$$l^{\mu} = \begin{bmatrix} 0 & -1 & 0 & 0 \end{bmatrix},$$

$$n^{\mu} = \begin{bmatrix} 1 & g_{00}/2 & 0 & 0 \end{bmatrix},$$

$$m^{\mu} = \frac{1}{\sqrt{2}r} \begin{bmatrix} 0 & r^{2}f & -1 & \frac{-i}{\sin\theta} \end{bmatrix},$$
 (9)

$$\overline{m}^{\mu} = \frac{1}{\sqrt{2}r} \begin{bmatrix} 0 & r^{2}f & -1 & \frac{i}{\sin\theta} \end{bmatrix}.$$

不难证明 (8)式和(9)式满足零标架的定义式(见文献(3]).

3.弯曲时空的 Dirac 场方程

在弯曲时空中的 Dirac 场方程为

$$(D + \varepsilon - \rho + ieA_{\mu}l^{\mu})F_{1}$$

$$+ (\bar{\delta} + \pi - \alpha + ieA_{\mu}\bar{m}^{\mu})F_{2}$$

$$= i\mu_{0}G_{1}/\sqrt{2},$$

$$(\Delta + \mu - \gamma + ieA_{\mu}n^{\mu})F_{2}$$

$$+ (\delta + \beta - \tau + ieA_{\mu}m^{\mu})F_{1}$$

$$= i\mu_{0}G_{2}/\sqrt{2},$$

$$(D + \varepsilon^{*} - \rho^{*} + ieA_{\mu}l^{\mu})G_{2}$$

$$- (\delta + \pi^{*} - \alpha^{*} + A_{\mu}m^{\mu})G_{1}$$

$$= i\mu_{0}F_{2}/\sqrt{2},$$

$$(\Delta + \mu^{*} - \gamma^{*} + ieA_{\mu}n^{\mu})G_{1}$$

$$- (\bar{\delta} + \beta^{*} - \tau^{*} + ieA_{\mu}\bar{m}^{\mu})G_{2}$$

$$= i\mu_{0}F_{1}/\sqrt{2},$$
(10)

式中

$$D = \partial_{00} = l^{\mu} \partial_{\mu} ,$$

$$\Delta = \partial_{11} = n^{\mu} \partial_{\mu} ,$$

$$\delta = \partial_{01} = m^{\mu} \partial_{\mu} ,$$

$$\bar{\delta} = \partial_{10} = \bar{m}^{\mu} \partial_{\mu} .$$
(11)

 μ_0 和 *e* 分别为 Dirac 粒子的静止质量和静止电量. A_{μ} 为黑洞所带电量所产生的电磁四矢 , F_1 , F_2 , G_1 和 G_2 为波矢量的四个分量 ,它们都是时空坐标 (*v* ,*r* , θ , φ)的函数.按照文献 3 提供的方法计算出 非零的旋系数如下:

$$\rho = \frac{1}{r} ,$$

$$\gamma = -\frac{1}{4} g_{00,1} ,$$

$$\beta = -\frac{1}{\sqrt{2}r} \cot\theta ,$$

$$\alpha = \frac{1}{2\sqrt{2}r} \cot\theta - \frac{f}{\sqrt{2}} ,$$

$$\mu = \frac{1}{2r} g_{00} ,$$

$$\tau = -\pi = -\frac{f}{\sqrt{2}} ,$$

$$\nu = -\frac{1}{\sqrt{2}} \left[\left(g_{00} - \frac{rg_{00,1}}{2} \right) f + r f_{,v} + \frac{g_{00,2}}{2r} \right] ,$$

式中

$$g_{00,1} = \partial g_{00} / \partial r ,$$

$$g_{00,2} = \partial g_{00} / \partial \theta ,$$

$$f_{,v} = \partial f / \partial v .$$

将(9)(11)和(12)式代入(10)式得到 Dirac 粒子的 一阶方程:

$$\sqrt{2}D_{1}F_{1} - (rfD_{2} - L_{-})F_{2}$$

$$= -i\mu_{0}G_{1}, \qquad (13)$$

$$\sqrt{2}\left(D_{0} + \frac{g_{00}}{2}D_{1} + \frac{g_{00,1}}{4}\right)F_{2} + (rfD_{1} - L_{+})F_{1}$$

$$= i\mu_{0}G_{2}, \qquad (14)$$

$$\sqrt{2D_1 G_2} + (rfD_2 - L_+)G_1$$

= $-i\mu_0 F_2$, (15)

$$\sqrt{2} \left(D_0 + \frac{g_{00}}{2} D_1 + \frac{g_{00,1}}{4} \right) G_1 - (rfD_1 - L_-)G_2$$

= $i\mu_0 F_1$, (16)

式中

$$D_{0} = \frac{\partial}{\partial v} + ieA_{0} ,$$

$$D_{1} = \frac{\partial}{\partial r} + ieA_{1} + \frac{1}{r} ,$$

$$D_{2} = \frac{\partial}{\partial r} + ieA_{1} + \frac{2}{r}$$

$$L_{\pm} = \frac{1}{r} \left(\frac{\partial}{\partial \theta} + ieA_{2} + \frac{1}{2} \cot \theta \right)$$

$$\pm \frac{i}{r \sin \theta} \left(\frac{\partial}{\partial \varphi} + ieA_{3} \right) .$$
(17)

为了考察 Dirac 粒子的辐射,现将(13)式中的 G₁与 (14)式中的 G₂分别代入(15)式和(16)式,得到 Dirac 粒子的二阶方程:

$$2\left(D_{0} + \frac{g_{00}}{2}D_{1} + \frac{g_{00,1}}{4}\right)D_{1}F_{1}$$

$$-\sqrt{2}\left(D_{0} + \frac{g_{00}}{2}D_{1} + \frac{g_{00,1}}{4}\right)\left(rfD_{2} - L_{-}\right)F_{2}$$

$$+\sqrt{2}\left(rfD_{1} - L_{-}\right)\left(D_{0} + \frac{g_{00}}{2}D_{1} + \frac{g_{00,1}}{4}\right)F_{2}$$

$$+\left(rfD_{1} - L_{-}\right)\left(rfD_{1} - L_{+}\right)F_{1} = \mu_{0}^{2}F_{1}, \quad (18)$$

$$2D_{1}\left(D_{0} + \frac{g_{00}}{2}D_{1} + \frac{g_{00,1}}{4}\right)F_{2}$$

$$+\sqrt{2}D_{1}\left(rfD_{1} - L_{+}\right)F_{1} - \sqrt{2}\left(rfD_{2} - L_{+}\right)D_{1}F_{1}$$

+($rfD_2 - L_+$)($rfD_2 - L_-$) $F_2 = \mu_0^2 F_2$. (19) 将(17)式代入到方程(18)和(19)中,经过计算整理 得到二阶方程的明显表达式:

$$\left(g_{00} + r^{2}f^{2}\right)\frac{\partial^{2}F_{1}}{\partial r^{2}} + 2\frac{\partial^{2}F_{1}}{\partial r\partial v} - 2f\frac{\partial^{2}F_{1}}{\partial r\partial \theta} + \frac{1}{r^{2}}\frac{\partial^{2}F_{1}}{\partial \theta^{2}} + \frac{1}{r^{2}\sin^{2}\theta}\frac{\partial^{2}F_{1}}{\partial \varphi^{2}} + \left[\frac{2g_{00}}{r} + \frac{g_{00,1}}{2} + 3rf^{2} + 2a\cos\theta + ie\mathcal{X}A_{0} + g_{00}A_{1} + r^{2}f^{2}A_{1} - fA_{2}\right]\frac{\partial F_{1}}{\partial r}$$

$$+ \left(ie2A_{1} + \frac{2}{r}\right)\frac{\partial F_{1}}{\partial v} - \left(ie2fA_{1} - \frac{ie2A_{2}}{r^{2}}\right)$$

$$+ \frac{f}{r} - \frac{1}{r^{2}}\cot\theta \frac{\partial F_{1}}{\partial \theta}$$

$$+ \frac{i}{r\sin\theta}\left(f + \frac{2eA_{3}}{r\sin\theta} - \frac{\cos\theta}{r\sin\theta}\right)\frac{\partial F_{1}}{\partial \varphi}$$

$$+ N_{11}F_{1} - \sqrt{2}\left[\left(fg_{00} - \frac{rfg_{00,1}}{2} + rf_{v} + \frac{g_{00,2}}{2r}\right)\frac{\partial F_{2}}{\partial r}\right]$$

$$+ f\frac{\partial F_{2}}{\partial v} + \frac{g_{00}}{2r^{2}}\frac{\partial F_{2}}{\partial \theta} - \frac{ig_{00}}{2r^{2}}\frac{\partial F_{2}}{\partial \varphi} + N_{12}F_{2}\right]$$

$$= \mu_{0}^{2}F_{1}, \qquad (20)$$

和

$$\left(g_{00} + r^{2}f^{2}\right)\frac{\partial^{2}F_{2}}{\partial r^{2}} + 2\frac{\partial^{2}F_{2}}{\partial r\partial v} - 2f\frac{\partial^{2}F_{2}}{\partial r\partial\theta}$$

$$+ \frac{1}{r^{2}}\frac{\partial^{2}F_{2}}{\partial\theta^{2}} + \frac{1}{r^{2}\sin^{2}\theta}\frac{\partial^{2}F_{2}}{\partial\varphi^{2}}$$

$$+ \left[\frac{2g_{00}}{r} + \frac{3g_{00,1}}{2} + 5rf^{2} + 2a\cos\theta$$

$$+ ie\chi A_{0} + g_{00}A_{1} + r^{2}f^{2}A_{1} - fA_{2}\right]\frac{\partial F_{2}}{\partial r}$$

$$+ \left(ie2A_{1} + \frac{2}{r}\right)\frac{\partial F_{2}}{\partial v} - \left(ie2fA_{1} - \frac{ie2A_{2}}{r^{2}}\right)$$

$$+ 3\frac{f}{r} - \frac{1}{r^{2}}\cot\theta\frac{\partial F_{2}}{\partial\theta}$$

$$+ \frac{i}{r\sin\theta}\left(\frac{2eA_{3}}{r\sin\theta} + \frac{\cos\theta}{r\sin\theta} - f\right)\frac{\partial F_{2}}{\partial\varphi}$$

$$+ N_{22}F_{2} + \sqrt{2}\left(\frac{1}{r^{2}}\frac{\partial F_{1}}{\partial\theta} + \frac{i}{r^{2}}\frac{\partial F_{1}}{\partial\varphi} + N_{21}\right)F_{1}$$

$$= \mu_{0}^{2}F_{2}. \qquad (21)$$

在(20)和(21)式中含有一阶导数和零阶导数交叉 项 特别的是在(20)式中交叉项 $\partial F_2/\partial r$ 的系数与旋 系数 \langle 见(12)式)成正比.另外 N_{11} , N_{21} 是 F_1 的系 数 $\langle N_{12}$, N_{22} 是 F_2 的系数 ,后面将看到该系数对 Dirac 粒子的热辐射没有关系,所以可以不予明显 表式.

4. 乌龟坐标变换

作乌龟坐标变换 ,令

$$r_{*} = \frac{1}{2\kappa(v_{0}, \theta_{0})} \ln[r - r_{H}(v, \theta)],$$

$$v_{*} = v - v_{0},$$

$$\theta_{*} = \theta - \theta_{0}.$$
(22)

式中,_{*r_H*为黑洞的事件视界; κ 为调节参数(后面将 看到 κ 即为表征黑洞 Hawking 辐射的温度函数)并} 且在乌龟坐标变换下不变; v_0 , θ_0 是与乌龟坐标变换无关的任意常数.

虽然方程 20)和(21)式是不能对 v ,r , θ , φ 经 过分离变量法面完全退耦的 ,而考察事件视界处 r_{H} 在时刻 v_0 、极角 θ_0 时的 Hawking 辐射效应 ,所以只 对它们在视界 $r_{H}(v_0, \theta_0)$ 处附近行为感兴趣.首先 将乌龟坐标变换(22)式代入方程(13)和(14)式 ,并 逐项乘以 $2\kappa(r - r_H)$ 后 ,再求其 $r \rightarrow r_H$ (表示 $v \rightarrow v_0$, $\theta \rightarrow \theta_0$)时的极限 ,得到在事件视界附近的 F_1 与 F_2 之间关系式

$$\sqrt{2} \frac{\partial F_1}{\partial r_*} - \left(rf + \frac{r_{H\theta}}{r} \right) \frac{\partial F_2}{\partial r_*} = 0 , \qquad (23)$$

$$\left(rf + \frac{r_{H\theta}}{r}\right)\frac{\partial F_1}{\partial r_*} + \sqrt{2}\left(\frac{g_{00}}{2} - r_{H\nu}\right)\frac{\partial F_2}{\partial r_*} = 0 \text{ (24)}$$

再将乌龟坐标变换(22)式代入到方程(20)和(21) 式 同样逐项乘以 2κ(r - r_H) 经整理后得到

$$A \frac{\partial^{2} F_{1}}{\partial r_{*}^{2}} + 2 \frac{\partial^{2} F_{1}}{\partial r_{*} \partial v_{*}} - 2 \left(f + \frac{r_{HP}}{r^{2}} \right) \frac{\partial^{2} F_{1}}{\partial r_{*} \partial \theta_{*}}$$

$$+ \left\{ - 2\kappa A + \frac{2g_{00}}{r} + \frac{g_{00,1}}{2} + 3rf^{2} + 2a\cos\theta - \frac{2r_{Hh}}{r} + \frac{fr_{HP}}{r} - \frac{r_{HP}}{r^{2}}\cot\theta - \frac{r_{HP}}{r^{2}} \right\}$$

$$+ ie2 \left[A_{0} + \left(g_{00} + r^{2}f^{2} - r_{Hh} + fr_{HP} \right) A_{1} - \left(f + \frac{r_{HP}}{r^{2}} \right) A_{2} \right] \right\} \frac{\partial F_{1}}{\partial r_{*}} + 2\kappa (r - r_{H})$$

$$\times \left[\frac{1}{r^{2}} \frac{\partial^{2} F_{1}}{\partial \theta_{*}^{2}} + \frac{1}{r^{2}\sin^{2}\theta} \frac{\partial^{2} F_{1}}{\partial \varphi_{*}^{2}} + \left(ie2A_{1} + \frac{2}{r} \right) \frac{\partial F_{1}}{\partial v_{*}} \right]$$

$$- \left(ie2fA_{1} - \frac{ie2A_{2}}{r^{2}} + \frac{f}{r} - \frac{1}{r^{2}}\cot\theta \right) \frac{\partial F_{1}}{\partial \theta_{*}}$$

$$+ \frac{i}{r\sin\theta} \left(f + \frac{2eA_{3}}{r\sin\theta} - \frac{\cos\theta}{r\sin\theta} \right) \frac{\partial F_{1}}{\partial \varphi}$$

$$+ \left(N_{11} - \mu_{0}^{2} \right) F_{1} + \sqrt{2}f \frac{\partial F_{2}}{\partial v_{*}} + \frac{g_{00}}{\sqrt{2}r^{2}} \frac{\partial F_{2}}{\partial \theta_{*}}$$

$$- \frac{ig_{00}}{\sqrt{2}r^{2}\sin\theta} \frac{\partial F_{2}}{\partial \varphi} + \sqrt{2}N_{12}F_{2} \right]$$

$$= \sqrt{2} \left(fg_{00} - \frac{rfg_{00,1}}{2} + rf_{*}r + \frac{g_{00,2}}{2r} \right)$$

$$(25)$$

和

$$\begin{split} A & \frac{\partial^2 F_2}{\partial r_*^2} + 2 \frac{\partial^2 F_2}{\partial r_* \partial v_*} - 2 \left(f + \frac{r_{H\theta}}{r^2} \right) \frac{\partial^2 F_2}{\partial r_* \partial \theta_*} \\ &+ \left\{ -2\kappa A + \frac{2g_{00}}{r} + \frac{3g_{00,l}}{2} + 5rf^2 + 2a\cos\theta \right. \end{split}$$

$$-\frac{2r_{H\nu}}{r} + \frac{3fr_{H\theta}}{r} - \frac{r_{H\theta}}{r^{2}}\cot\theta - \frac{r_{H\theta}}{r^{2}}$$

$$+ ie2\Big[A_{0} + (g_{00} + r^{2}f^{2} - r_{H\nu} + fr_{H\theta})A_{1}$$

$$- \left(f + \frac{r_{H\theta}}{r^{2}}\right)A_{2}\Big]\Big\}\frac{\partial F_{2}}{\partial r_{*}} + 2\kappa(r - r_{H})$$

$$\times \Big[\frac{1}{r^{2}}\frac{\partial^{2}F_{2}}{\partial \theta_{*}^{2}} + \frac{1}{r^{2}\sin^{2}\theta}\frac{\partial^{2}F_{2}}{\partial \varphi^{2}} + \left(ie2A_{1} + \frac{2}{r}\right)\frac{\partial F_{2}}{\partial v_{*}}$$

$$- \left(ie2fA_{1} - \frac{ie2A_{2}}{r^{2}} + \frac{3f}{r} - \frac{1}{r^{2}}\cot\theta\right)\frac{\partial F_{2}}{\partial \theta_{*}}$$

$$+ \frac{i}{r\sin\theta}\Big(\frac{\cos\theta}{r\sin\theta} + \frac{2eA_{3}}{r\sin\theta} - f\Big)\frac{\partial F_{2}}{\partial \varphi}$$

$$+ (N_{22} - \mu_{0}^{2})F_{2} + \frac{\sqrt{2}}{r^{2}}\frac{\partial F_{1}}{\partial \theta_{*}}$$

$$+ \frac{i\sqrt{2}}{r^{2}\sin\theta}\frac{\partial F_{1}}{\partial \varphi} + \sqrt{2}N_{21}F_{1}\Big]$$

$$= \frac{\sqrt{2}r_{H\theta}}{r^{2}}\frac{\partial F_{1}}{\partial r_{*}}, \qquad (26)$$

$$\vec{x} \oplus r_{H\theta\theta} = \partial^{2}r_{H}/\partial\theta^{2}|_{r=r_{H}}, \vec{m}, \vec{x} \otimes A, \vec{x}$$

$$A = \frac{g_{00} + r^{2}f^{2} - 2r_{H\nu} + 2fr_{H\theta} + r_{H\theta}^{2}/r^{2}}{2\kappa(r - r_{H})}. \qquad (27)$$

5. 视界面附近的标准波动方程

我们希望能将用乌龟坐标表示的 Dirac 二阶方 程,在黑洞事件视界面附近化为单一分量的标准波 动方程,要求当 $r \rightarrow r_H$ 时,方程(25)和(26)式中第一 项 $\partial^2 F / \partial r_*^2$ 系数 A 的极限应为一个常数,必须有其 分子在 $r \rightarrow r_H$ 时的极限趋近于零,得到确定黑洞事 件视界位置的方程:

$$1 - \frac{2M}{r_{H}} - 2ar_{H}\cos\theta + \frac{Q^{2}}{r_{H}^{2}} - 4a\frac{Q^{2}}{r_{H}}\cos\theta$$
$$- 2r_{Hv} + 2fr_{H\theta} + \frac{r_{H\theta}^{2}}{r_{H}^{2}} = 0, \qquad (28)$$

与从零曲面方程导出的结果(见(7)式)一致,但两者 成立的范围不同.方程(7)式对任意时刻 v,任意极 角 θ 成立;而方程(28)式只对固定时刻 v_0 ,固定极 角 θ_0 成立,是方程(7)式在 $v = v_0$, $\theta = \theta_0$ 时的特殊 情形.在视界面附近, $\partial^2 F/\partial r_*^2$ 系数 A 的极限值为 0/0型,可用罗必塔法则求极限并调节参数 κ ,使该 系数的极限趋近于 1,得到

$$\kappa = (M + 2aQ^2\cos\theta)/r^2 - (Q^2 + r_{H^2}^2)/r^2 - a\cos\theta \mid_{r=r_H}.$$
 (29)

我们看出,在视界面附近二阶方程(25)和(26)式中

还含有一阶导数的交叉项,它使得每一个分量所满 足的方程没有完全化为单一分量的标准波动方程. 我们按吴双清和蔡勖建议的方法,将关系式(23)式 分别代入(25)和(26)式中,每一个方程中的一阶导 数交叉项就被完全消除,可得到

$$\frac{\partial^{2} F_{1}}{\partial r_{*}^{2}} + 2 \frac{\partial^{2} F_{1}}{\partial r_{*} \partial v_{*}} - 2 \left(f + \frac{r_{H\theta}}{r^{2}} \right) \frac{\partial^{2} F_{1}}{\partial r_{*} \partial \theta_{*}} + \left\{ \frac{g_{00}}{r} - \frac{g_{00,1}}{2} + 2a\cos\theta - \frac{fr_{H\theta}}{r} \right. + \frac{r_{H\theta}^{2}}{r^{3}} - \frac{r_{H\theta}}{r^{2}}\cot\theta - \frac{r_{H\theta\theta}}{r^{2}} + ie2 \left[A_{0} + \left(g_{00} + r^{2}f^{2} - r_{Hv} + fr_{H\theta} \right) A_{1} - \left(f + \frac{r_{H\theta}}{r^{2}} \right) A_{2} \right] \right\} \frac{\partial F_{1}}{\partial r_{*}} = \frac{2r}{r^{2}f + r_{H\theta}} \left(2fg_{00} - rfg_{00,1} + 2rf_{*v} + \frac{g_{00,2}}{r^{2}} + 2fr_{Hv} + \frac{g_{00}r_{H\theta}}{r^{2}} \right) \frac{\partial F_{1}}{\partial r_{*}} , \qquad (30)$$

和

$$\frac{\partial^2 F_2}{\partial r_*^2} + 2 \frac{\partial^2 F_2}{\partial r_* \partial v_*} - 2 \left(f + \frac{r_{HI}}{r^2} \right) \frac{\partial^2 F_2}{\partial r_* \partial \theta_*} + \left\{ \frac{g_{00}}{r} + \frac{g_{00,I}}{2} + 2rf^2 + 2a\cos\theta \right. + \frac{fr_{HI}}{r} + \frac{r_{HI}^2}{r^3} - \frac{r_{HI}}{r^2}\cot\theta - \frac{r_{HII}}{r^2} + ie2 \left[A_0 + \left(g_{00} + r^2 f^2 - r_{HI} + fr_{HII} \right) A_1 \right] - \left(f + \frac{r_{HII}}{r^2} \right) A_2 \right] \frac{\partial F_2}{\partial r_*} = \frac{\left(r^2 f + r_{HII} \right) r_{HII}}{r^3} \frac{\partial F_2}{\partial r_*} , \qquad (31)$$

在视界面附近的标准波动方程可统一写成

$$\frac{\partial^{2} F}{\partial r_{*}^{2}} + 2 \frac{\partial^{2} F}{\partial r_{*} \partial v_{*}} + 2B \frac{\partial^{2} F}{\partial r_{*} \partial \theta_{*}}$$
$$+ (C_{j} + i2\omega_{0}) \frac{\partial F}{\partial r_{*}} = 0, \qquad (32)$$

式中各项的系数为

$$B = -\left(f + \frac{r_{H\theta}}{r^2}\right)\Big|_{r=r_H},$$

$$\omega_0 = e\left[A_0 + \left(g_{00} + r^2 f^2 - r_{Hv} + fr_{H\theta}\right)A_1 - \left(f + \frac{r_{H\theta}}{r^2}\right)A_2\right]\Big|_{r=r_H},$$

$$C_1 = \frac{g_{00}}{r} - \frac{g_{00,1}}{2} + 2a\cos\theta$$

$$+ \frac{r_{H\theta}}{r} \left(\frac{r_{H\theta}}{r^2} - f - \frac{\cot\theta}{r} \right) - \frac{r_{H\theta\theta}}{r^2}$$
$$- \frac{2r}{r^2 f + r_{H\theta}} \left[\left(2g_{00} + 2r_{Hv} - rg_{00,1} \right) f + 2rf_{,v} + \frac{g_{00,2}}{r} + \frac{g_{00}r_{H\theta}}{r^2} \right] \Big|_{r=r_H},$$
$$C_2 = \frac{g_{00}}{r} + \frac{g_{00,1}}{2} + 2rf^2 + 2a\cos\theta$$
$$- \frac{r_{H\theta}}{r^2} \cot\theta - \frac{r_{H\theta\theta}}{r^2} \Big|_{r=r_H}.$$
(33)

6. 解析延拓及辐射谱

在视界面附近,方程(32)的解可写成

$$F = R(r_*)\Theta(\theta_*)e^{-i\omega v_* + il\varphi}.$$
 (34)

将(34) 式代入(32) 式可得

$$\frac{\mathrm{d}\Theta}{\mathrm{d}\theta} = \lambda\Theta ,$$

$$\frac{\mathrm{d}^{2}R(r_{*})}{\mathrm{d}r_{*}^{2}} + [C_{j} + 2B\lambda - i\mathcal{X} \omega - \omega_{0}]\frac{\mathrm{d}R(r_{*})}{\mathrm{d}r_{*}} = 0 ,$$
(35)

入射波与出射波分别为

$$F^{\text{in}} = e^{-i\omega v_{*} + \lambda \theta_{*} + il\varphi}$$
,
 $F^{\text{out}} = e^{-i\omega v_{*} + \lambda \theta_{*} + il\varphi} e^{-(C_{j} + 2B\lambda)r_{*} + i\chi\omega - \omega_{0})r_{*}}$. (36)

在视界面附近出射波可写成

$$F^{\text{out}} = e^{-i\omega v_{*} + \lambda \theta_{*} + il\varphi} (r - r_{H})^{-(C_{j} + 2B\lambda)2\kappa} (r - r_{H})^{(\omega - \omega_{0})\kappa}.$$
(37)

显然,F^{out}在视界面处具有非解析(36)式只能 描述视界面处外面的出射粒子,不能描述视界面处 内部的出射粒子,为此通过下半复平面绕过视界面 外解析延拓到视界面内部,即

$$r - r_{H} \rightarrow | r - r_{H} | e^{-i\pi} = (r_{H} - r)e^{-i\pi},$$

于是 在视界面内部的出射波为

$$\tilde{F}^{\text{out}} = e^{-i\omega v_{*} + \lambda \theta_{*} + il\varphi} [(r_{H} - r)e^{-i\pi}]^{(C_{j} + 2B\lambda)2\kappa}$$

$$\times [(r_{H} - r)e^{-i\pi}]^{\omega - \omega_{0}} \gamma_{\kappa}$$

$$= e^{-i\omega v_{*} + \lambda \theta_{*} + il\varphi} e^{-(C_{j} + 2B\lambda)r_{*} + i(\omega - \omega_{0})r_{*}}$$

$$\times e^{i\pi (C_{j} + 2B\lambda)^{2\kappa}} e^{\pi (\omega - \omega_{0}) \gamma_{\kappa}}.$$
(38)

出射波在视界面处的散射概率为

$$F^{\text{out}}/\tilde{F}^{\text{out}}|^2 = e^{-2\pi (\omega - \omega_0) \gamma_\kappa}$$
, (39)

根据 Damour-Ruffinf¹⁴¹和 Sannan^[15]方法,不难得到出 射波的黑体谱为

$$N_{\omega}^{2} = \left[e^{(\omega - \omega_{0}) \gamma_{k_{B}} T} + 1 \right]^{-1}.$$
 (40)

其中 $k_{\rm B}$ 为玻尔兹曼常数 , $T = \kappa/2\pi k_{\rm B}$ 为黑洞辐射温

度.κ 是由(29)式给出,从这里可以看出 κ 正是 Hawking 辐射的温度函数.

7. 结 论

总之,求解直线加速 Kinnersley 黑洞中 Dirac 方 程,将耦合 Dirac 方程中每一个方程式化成两个分量 的耦合方程,利用新的乌龟坐标变换,得到决定事件 视界的位置和辐射温度函数,将两个分量的耦合方 程进行了退耦,使得每一个分量满足的二阶方程化 成了标准的波动方程,证明了在直线加速 Kinnersley 黑洞中荷电 Dirac 粒子也具有 Hawking 辐射谱.结果 表明事件视界的位置和辐射温度不仅随时间变化, 而且依赖于角度.

如果利用旧的乌龟坐标变换来讨论该问题,得 到的辐射温度函数 κ 为

$$\kappa = \frac{1}{r} \frac{Mr - ar^3 \cos\theta - Q^2 + 2aQ^2 r \cos\theta - r_{H\theta}^2}{2Mr + 2ar^3 \cos\theta - Q^2 + 4aQ^2 r \cos\theta + r_{H\theta}^2} \Big|_{r=r_H}$$
(41)

与之相比 (29) 式所表示的比较简洁.在不同的乌龟 坐标变换下,得到的视界面位置方程和辐射谱的形 式是完全相同的,但是辐射温度函数有些差别.

作者曾与赵峥教授和朱建阳教授进行过讨论,在此表示 感谢.

- [1] Yang B, Zhao Z 1994 Acta Phys. Sin. 43 858(in Chinese I 杨 波、赵 峥 1994 物理学报 43 858]
- [2] Cao J L, Peng F Z 1998 Acta Phys. Sin. 47 177 (in Chinese] 曹 江陵、彭方志 1998 物理学报 47 177]
- [3] Zhao Z 1999 Thermal Properties of Black Hole and Singularities of Space-time (Beijing: Beijing Normal University Press) (in Chinese)
 [赵 峥 1999 黑洞的热性质与时空奇异性(北京:北京师范 大学出版社)]
- [4] Wu S Q 2002 Doctor Dissertation (Wuhan: Huazhong Normal University) in Chinese [吴双清 2002 博士学位论文(武汉:华 中师范大学)]
- [5] Wu S Q , Cai X 2002 Chin. Phys. 11 661
- [6] Wu S Q , Cai X 2002 Chin . Phys . Lett . 19 141
- [7] Wu S Q, Cai X 2002 Int. J. Theor. Phys. 41 641

- [8] Wu S Q , Cai X 2002 Gen . Rel . Grav . 34 1207
- [9] Wu S Q, Zeng Y, Cai X, Yan M L 2003 Acta Phys. Sin. 52 1340
 (in Chinese)[吴双清、曾 瑜、蔡 勖、闫沐霖 2003 物理学报 52 1340]
- [10] Wu S Q , Yan M L 2003 Chin . Phys . Lett . 20 1913
- [11] Zhang J Y, Zhao Z 2003 Acta Phys. Sin. 52 2096 in Chinese)[张 靖仪、赵 峥 2003 物理学报 52 2096]
- [12] Cao J L 2006 Acta Phys. Sin. 55 2682(in Chinese] 曹江陵 2006 物理学报 55 2682]
- [13] Niu Z F, Liu W B 2005 Acta Phys. Sin. 54 475(in Chinese] 牛振风、刘文彪 2005 物理学报 54 475]
- [14] Damour T, Ruffini R 1976 Phys. Rev. D 14 332
- [15] Sannan S 1988 Gen. Rel. Grav. 20 239

Hawking radiation of Dirac particles in a rectilinearly accelerating Kinnersley black hole *

Yang Bo

(Physics and Electronic Engineering , College of Chongqing Three Gorgee University , Chongqing 404000 , China)
 (Received 19 April 2007 ; revised manuscript received 14 June 2007)

Abstract

In a uniformly accelerating space time, the coupling Dirac equation can be simplified to the second-order equation. Using new tortoise coordinates, we further simplify the equation to the standard wave-equation near the horizon and get the temperature function of radiation and the Hawking radiation spectrum.

Keywords: black hole, Dirac equation, tortoise coordinate transformation, Hawking radiation **PACC**: 9760L, 0420

^{*} Project supported by the Scientific Technique Research Event Program of Chongqing Municipal Education Commission of China (Grant No. KJ071111).