

匀加速直线运动的 Kerr 黑洞的特征曲面^{*}

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研究了作匀加速直线运动的变质量 Kerr 黑洞的三类特征曲面, 即事件视界、表观视界和类时极限面(无限红移面). 指出在变质量(存在吸积或蒸发)情况下, 这类黑洞的三类特征曲面不再重合.

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1 引 言

近年来, 黑洞的反作用已被大量地研究^[1-3], 在文献[4]中研究了动态 Kerr 黑洞的类时极限面, 表观视界和事件视界的位置, 它们不再重合. 在动态情况下, 黑洞外能层变厚, 内能层变薄. 文献[5]中在 de Sitter 膨胀宇宙中研究了一个 Vaidya 蒸发黑洞模型, 讨论了表观视界和事件视界, 发现它们不再重合. 文献[6]中用 C 度规研究了作匀加速且旋转的稳态黑洞的量子热性质, 由零超曲面方程决定出事件视界的位置. 而本文进一步研究了作匀加速直线运动的动态 Kerr 黑洞的三类特征曲面.

匀加速直线运动的 Kerr 黑洞的线元用超前爱丁顿坐标表示为^[7]

$$\begin{aligned} ds^2 = & (1 - F)dv^2 - 2dvdr - 2fp^2dv d\theta + 2\sin^2\theta AFdv d\varphi \\ & + 2A\sin^2\theta dr d\varphi - p^2d\theta + 2Afp^2\sin^2\theta d\theta d\varphi \\ & - \sin^2\theta[A^2F\sin^2\theta + A^2 + r^2]d\varphi^2, \end{aligned} \quad (1)$$

其中

$$\begin{aligned} F & \equiv 2\arccos\theta + 2mr/p^2 + p^2f^2, \\ f & \equiv -a\sin\theta, \\ p^2 & \equiv r^2 + A^2\cos^2\theta. \end{aligned} \quad (2)$$

转动轴的北极 $\theta=0$, 指向加速度方向, A 为单位质量的角动量, 是一个常量; a 为加

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速度, 是一个常量; m 为源质量, 是 v 的函数.

2 类时极限面(TLS)

类时极限面的定义为

$$g_{vv} = \delta_{ab} \left(\frac{\partial}{\partial v} \right)^a \left(\frac{\partial}{\partial v} \right)^b = 0. \quad (3)$$

由方程(1)可知:

$$g_{vv} = 1 - F = 0, \quad (4)$$

即

$$-a^2 \sin^2 \theta r^4 - 2a \cos \theta r^3 + (1 - 2A^2 a^2 \sin^2 \theta \cos^2 \theta) r^2 - (2m + 2A^2 a \cos^3 \theta) r + A^2 \cos^2 \theta - A^4 a^2 \sin^2 \theta \cos^4 \theta = 0. \quad (5)$$

解方程(5)得4个根, 其中1个是负根无物理意义, 其余3个根为:

$$r_{\text{TLS}_1} = -D + E + R - \frac{1}{2} \frac{\cos \theta}{a \sin^2 \theta}, \quad (6)$$

$$r_{\text{TLS}_2} = D - E + R - \frac{1}{2} \frac{\cos \theta}{a \sin^2 \theta}, \quad (7)$$

$$r_{\text{TLS}_3} = D + E - R - \frac{1}{2} \frac{\cos \theta}{a \sin^2 \theta}, \quad (8)$$

$$\text{其中 } D = \frac{1}{\sqrt{6} a \sin \theta} [1 - a \cos(\varphi/3 - \pi/3)]^{1/2}, \quad (9)$$

$$E = \frac{1}{\sqrt{6} a \sin \theta} [1 - a \cos(\varphi/3 + \pi/3)]^{1/2}, \quad (10)$$

$$R = \frac{1}{\sqrt{6} a \sin \theta} [1 + a \cos(\varphi/3)]^{1/2}, \quad (11)$$

$$\alpha = \left[\left(1 - 2A^2 a^2 \sin^2 \theta \cos^2 \theta + \frac{3}{2} \cot^2 \theta \right)^2 - 12 \left(\frac{3}{16} \frac{\cos^4 \theta}{a^2 \sin^6 \theta} + \frac{1}{4} \frac{\cos^2 \theta}{a^2 \sin^4 \theta} + \frac{m \cos \theta}{a \sin^2 \theta} + \frac{1}{2} \frac{A^2 \cos^4 \theta}{\sin^2 \theta} + A^2 \cos^2 \theta - A^4 a^2 \sin^2 \theta \cos^4 \theta \right) \right]^{1/2}, \quad (12)$$

$$\varphi = \arccos(-\beta/\alpha^3). \quad (13)$$

当 $0 < \varphi < \pi$ 时,

$$\begin{aligned} \beta = & \left(1 - 2A^2 a^2 \sin^2 \theta \cos^2 \theta + \frac{3}{2} \cot^2 \theta \right)^3 + 30 a^2 \sin^2 \theta \left(1 - 2A^2 a^2 \sin^2 \theta \cos^2 \theta \right. \\ & + \frac{3}{2} \cot^2 \theta \left[\frac{3}{16} \frac{\cos^4 \theta}{a^2 \sin^6 \theta} + \frac{1}{4} \frac{\cos^2 \theta}{a^2 \sin^4 \theta} + \frac{m \cos \theta}{a \sin^2 \theta} + \frac{1}{2} \frac{A^2 \cos^4 \theta}{\sin^2 \theta} \right. \\ & \left. \left. + A^2 \cos^2 \theta - A^4 a^2 \sin^2 \theta \cos^4 \theta \right] - 27 a^2 \sin^2 \theta \left(2m + \frac{\cos \theta}{a \sin^2 \theta} + \frac{3}{2} \frac{\cos^3 \theta}{a \sin^4 \theta} \right)^2. \end{aligned} \quad (14)$$

r_{TLS_1} 为 Rindler 的类时极限面, r_{TLS_2} 和 r_{TLS_3} 分别为匀加速直线运动的 Kerr 黑洞的外部 and 内部类时极限面.

当 $A=0$, $m=0$ 时, 方程(5)化为

$$a^2 \sin^2 \theta r^4 + 2a \cos \theta r^3 - r^2 = 0, \quad (15)$$

解之得 Rindler 的类时极限面

$$r = [a(1 + \cos \theta)]^{-1}. \quad (16)$$

当 $a=0$ 时(即无加速静止不动的黑洞), 方程(5)变为

$$r^2 - 2mr + A^2 \cos^2 \theta = 0, \quad (17)$$

解之得

$$r_{\text{TLS}}^{\pm} = m \pm \sqrt{m^2 - A^2 \cos^2 \theta}. \quad (18)$$

此时的类时极限面与动态 Kerr 黑洞的类时极限面相一致.

对于蒸发黑洞 $\dot{m} = \frac{dm}{dv} < 0$, 则不难看出外类时极限面逐渐收缩, 而内类时极限面逐渐膨胀.

对于吸积黑洞 $\dot{m} = \frac{dm}{dv} > 0$, 则有相反的结果.

3 表观视界(AH)

表观视界定义为对于出射光子的最外陷获面(俘获面), 从它出发的两簇类光测地线汇的膨胀 Θ 为零.

由方程(1)可知其度规为

$$(g_{\mu\nu}) = \begin{pmatrix} 1-F & -1 & -fp^2 & AF\sin^2 \theta \\ -1 & 0 & 0 & A\sin^2 \theta \\ -fp^2 & 0 & -p^2 & Af p^2 \sin^2 \theta \\ AF\sin^2 \theta & A\sin^2 \theta & Af p^2 \sin^2 \theta & -\sin^2 \theta [A^2 F \sin^2 \theta + A^2 + r^2] \end{pmatrix}. \quad (19)$$

不难算出行列式和逆变度规

$$g = -p^4 \sin^2 \theta, \quad (20)$$

$$(g^{\mu\nu}) = \begin{pmatrix} -\frac{A^2}{p^2} \sin^2 \theta & -\left(1 + \frac{A^2}{p^2} \sin^2 \theta\right) & 0 & -\frac{A}{p^2} \\ -\left(1 + \frac{A^2}{p^2} \sin^2 \theta\right) & \left(F - 1 - f^2 p^2 - \frac{A^2}{p^2} \sin^2 \theta\right) & f & -\frac{A}{p^2} \\ 0 & f & -\frac{1}{p^2} & 0 \\ -\frac{A}{p^2} & -\frac{A}{p^2} & 0 & -\frac{1}{p^2 \sin^2 \theta} \end{pmatrix}. \quad (21)$$

可算出不为零的联络(见附录), 下面引入零标架. 先将方程(1)对称化得

$$\begin{aligned} ds^2 = & [\gamma(r^2 + A^2)(dv - A\sin^2 \theta d\varphi)] \left[\frac{1}{r^2 + A^2} \left(\frac{\Delta}{2} dv - \gamma^{-1} dr \right. \right. \\ & \left. \left. - f\gamma^{-2} d\theta - \frac{\Delta}{2} A\sin^2 \theta d\varphi \right) \right] + \left[\frac{1}{r^2 + A^2} \left(\frac{\Delta}{2} dv - \gamma^{-1} dr - f\gamma^{-2} d\theta \right. \right. \\ & \left. \left. - \frac{\Delta}{2} A\sin^2 \theta d\varphi \right) \right] [\gamma(r^2 + A^2)(dv - A\sin^2 \theta d\varphi)] - \sqrt{\frac{\gamma}{2}} [i A\sin \theta dv \\ & + \gamma^{-1} d\theta - i(r^2 + A^2)\sin \theta d\varphi] \sqrt{\frac{\gamma}{2}} [-i A\sin \theta dv + \gamma^{-1} d\theta \end{aligned}$$

$$\begin{aligned}
& + i(r^2 + A^2)\sin\theta d\varphi] - \sqrt{\frac{\gamma}{2}}[-iA\sin\theta dv + \gamma^{-1}d\theta + i(r^2 + A^2)\sin\theta d\varphi] \\
& \times \sqrt{\frac{\gamma}{2}}[iA\sin\theta dv + \gamma^{-1}d\theta - i(r^2 + A^2)\sin\theta d\varphi],
\end{aligned} \quad (22)$$

其中 $\gamma = 1/p^2$, $\Delta = r^2 + A^2 - \gamma^{-1}F$. (23)

得零标架为(不为零的联络见附录)

$$\begin{aligned}
n_\mu &= \gamma(r^2 + A^2)(1, 0, 0, -A\sin^2\theta), \\
l_\mu &= \frac{1}{r^2 + A^2}\left(\frac{\Delta}{2}, -\gamma^{-1}, -f\gamma^{-2}, -\frac{\Delta}{2}A\sin^2\theta\right), \\
m_\mu &= \sqrt{\frac{\gamma}{2}}(-iA\sin\theta, 0, \gamma^{-1}, i(r^2 + A^2)\sin\theta), \\
\bar{m}_\mu &= \sqrt{\frac{\gamma}{2}}(iA\sin\theta, 0, \gamma^{-1}, -i(r^2 + A^2)\sin\theta).
\end{aligned} \quad (24)$$

用度规升降指标得逆变零标架为

$$\begin{aligned}
n^\mu &= \gamma(r^2 + A^2)(0, -1, 0, 0), \\
l^\mu &= \left(1, \frac{\Delta}{2(r^2 + A^2)}, 0, \frac{A}{r^2 + A^2}\right), \\
m^\mu &= \sqrt{\frac{\gamma}{2}}\left(iA\sin\theta, 0, 1, \frac{i}{\sin\theta}\right), \\
\bar{m}^\mu &= \sqrt{\frac{\gamma}{2}}\left(-iA\sin\theta, 0, 1, -\frac{i}{\sin\theta}\right).
\end{aligned} \quad (25)$$

不难验证, (24), (25)式满足零标架条件

$$g_{\mu\nu} = l_\mu n_\nu + n_\mu l_\nu - m_\mu \bar{m}_\nu - \bar{m}_\mu m_\nu, \quad (26)$$

$$\begin{aligned}
n_\mu n^\mu &= l_\mu l^\mu = m_\mu m^\mu = \bar{m}_\mu \bar{m}^\mu = 0, \\
n^\mu l_\mu &= m_\mu \bar{m}^\mu = 1,
\end{aligned} \quad (27)$$

$$\begin{aligned}
n_\mu m^\mu &= n_\mu \bar{m}^\mu = l_\mu m^\mu = l_\mu \bar{m}^\mu = 0. \\
\Theta &= l^\mu_{;\mu} - \kappa,
\end{aligned} \quad (28)$$

而膨胀

其中 $\kappa = -\frac{1}{2}\frac{dF}{dr} + \frac{A^2\sin^2\theta}{2(r^2 + A^2)}\frac{dF}{dr} - \frac{A^2Fr\sin^2\theta}{(r^2 + A^2)^2}$, (29)

所以 $\Theta = \frac{r\Delta}{p^2(r^2 + A^2)}$. (30)

可见表观视界由 $\Delta=0$ 来定, 即

$$\begin{aligned}
& -a^2\sin^2\theta r^4 - 2a\cos\theta r^3 + (1 - 2A^2a^2\sin^2\theta\cos^2\theta)r^2 \\
& - (2m + 2A^2a\cos^3\theta)r - A^4a^2\sin^2\theta\cos^4\theta + A^2 = 0.
\end{aligned} \quad (31)$$

解方程(31)可得 4 个根, 其中 1 个是负根无物理意义, 其余 3 个根为

$$r_{\text{AH}_1} = -D' + E' + R' - \frac{1}{2}\frac{\cos\theta}{a\sin^2\theta}, \quad (32)$$

$$r_{\text{AH}_2} = D' - E' + R' - \frac{1}{2}\frac{\cos\theta}{a\sin^2\theta}, \quad (33)$$

$$r_{\text{AH}_3} = D' + E' - R' - \frac{1}{2} \frac{\cos \theta}{a \sin^2 \theta}, \quad (34)$$

$$\text{其中 } D' = \frac{1}{\sqrt{6} a \sin \theta} [1 - \alpha' \cos(\varphi'/3 - \pi/3)]^{1/2}, \quad (35)$$

$$E' = \frac{1}{\sqrt{6} a \sin \theta} [1 - \alpha' \cos(\varphi'/3 + \pi/3)]^{1/2}, \quad (36)$$

$$R' = \frac{1}{\sqrt{6} a \sin \theta} [1 + \alpha' \cos(\varphi'/3)]^{1/2}, \quad (37)$$

$$\alpha' = \left[\left(1 - 2A^2 a^2 \sin^2 \theta \cos^2 \theta + \frac{3}{2} \cot^2 \theta \right)^2 - 12 \left(\frac{3}{16} \frac{\cos^4 \theta}{a^2 \sin^6 \theta} + \frac{1}{4} \frac{\cos^2 \theta}{a^2 \sin^4 \theta} + \frac{m \cos \theta}{a \sin^2 \theta} + \frac{1}{2} \frac{A^2 \cos^4 \theta}{\sin^2 \theta} + A^2 - A^4 a^2 \sin^2 \theta \cos^4 \theta \right) \right]^{1/2}, \quad (38)$$

$$\varphi' = \arccos(-\beta'/\alpha'^3). \quad (39)$$

当 $0 < \varphi' < \pi$ 时,

$$\begin{aligned} \beta' = & \left(1 - 2A^2 a^2 \sin^2 \theta \cos^2 \theta + \frac{3}{2} \cot^2 \theta \right)^3 + 30 a^2 \sin^2 \theta \left(1 - 2A^2 a^2 \sin^2 \theta \cos^2 \theta \right. \\ & \left. + \frac{3}{2} \cot^2 \theta \right) \left[\frac{3}{16} \frac{\cos^4 \theta}{a^2 \sin^6 \theta} + \frac{1}{4} \frac{\cos^2 \theta}{a^2 \sin^4 \theta} + \frac{m \cos \theta}{a \sin^2 \theta} + \frac{1}{2} \frac{A^2 \cos^4 \theta}{\sin^2 \theta} \right. \\ & \left. + A^2 - A^4 a^2 \sin^2 \theta \cos^4 \theta \right] - 27 \left(2m + \frac{\cos \theta}{a \sin^2 \theta} + \frac{3}{2} \frac{\cos^3 \theta}{a \sin^4 \theta} \right)^2 a^2 \sin^2 \theta. \end{aligned} \quad (40)$$

r_{AH_1} 为 Rindler 的表观视界, r_{AH_2} 和 r_{AH_3} 分别为匀加速直线运动的 Kerr 黑洞的外部 and 内部的表观视界.

当 $A=0$, $m=0$ 时得 Rindler 表观视界, 方程(31)变为

$$a^2 \sin^2 \theta r^4 + 2 a \cos \theta r^3 - r^2 = 0, \quad (41)$$

解之得

$$r = [a(1 + \cos \theta)]^{-1}. \quad (42)$$

当 $a=0$ 时, 则表观视界由 $\Delta = r^2 + A^2 - 2mr = 0$ 来确定, 此时

$$r_{\text{AH}}^{\pm} = m \pm \sqrt{m^2 - A^2}. \quad (43)$$

也就是说此时的表观视界与不作加速运动的 Kerr 黑洞的表观视界相一致, 即回到 Kerr 黑洞.

4 事件视界(EH)

确定动态时空中事件视界的位置是一个非常困难的问题. York^[8]认为, 事件视界是光子的黏着面, 不允许光子从这里逃向远方, 因此我们由

$$\frac{d}{dv} = l^a \nabla_a = \frac{\partial}{\partial v} + \frac{\Delta}{2(r^2 + A^2)} \frac{\partial}{\partial r} + \frac{A}{r^2 + A^2} \frac{\partial}{\partial \varphi} \quad (44)$$

可得

$$\dot{r} = \frac{dr}{dv} = \frac{\Delta}{2(r^2 + A^2)}, \quad (45)$$

$$\ddot{r} = -\frac{p^2}{2(r^2 + A^2)} \frac{dF}{dv} + \dot{r} \left(\frac{r(1-F)}{r^2 + A^2} - \frac{p^2}{2(r^2 + A^2)} \frac{dF}{dr} - \frac{2r\dot{r}}{r^2 + A^2} \right). \quad (46)$$

由(45)和(46)式可知,当 $\frac{dF}{dv} < 0$ 时,在外表观视界处即 r_{AH}^+ 处

$$\dot{r} = 0, \quad \ddot{r} > 0. \quad (47)$$

因此,光子可以从表观视界逃到远方,这不符合事件视界的定义.可见,对于匀加速直线运动的动态 Kerr 黑洞而言,表观视界与事件视界不再重合,根据 York 的“工作定义”,即当把辐射量 $L = -\dot{m}$ 看作小量时,在一级近似下,1) 事件视界应严格类光;2) 光子应长时间滞留在事件视界上.这就是说,在 $O(L)$ 的数量级,事件视界应由

$$\dot{r}_{EH} \approx 0, \quad \ddot{r}_{EH} \approx 0 \quad (48)$$

决定,利用叠代法,先令(46)式中的 $\ddot{r} = 0$ 得

$$-\frac{p^2}{2(r^2 + A^2)} \frac{dF}{dv} + \dot{r} \left(\frac{r(1-F)}{r^2 + A^2} - \frac{p^2}{2(r^2 + A^2)} \frac{dF}{dr} - \frac{2r\dot{r}}{r^2 + A^2} \right) = 0. \quad (49)$$

当 $\frac{dF}{dv} = 0$ 时,即稳态情况下,

$$\dot{r} \left(\frac{r(1-F)}{r^2 + A^2} - \frac{p^2}{2(r^2 + A^2)} \frac{dF}{dr} - \frac{2r\dot{r}}{r^2 + A^2} \right) = 0. \quad (50)$$

解(50)式得

$$\dot{r} = \frac{1}{2} \left(1 - 2 \arccos \theta - \frac{2mr}{p^2} - p^2 f^2 \right) - \frac{p^2}{4} \left(2 \arccos \theta + 2rf^2 + \frac{2mp^2 - 4mr}{p^4} \right). \quad (51)$$

这时的 r 应取 r_{AH}^+ ,则

$$\begin{aligned} \dot{r} = & \frac{1}{2} \left[1 - 2 \arccos \theta - \frac{2mr_{AH}^+}{r_{AH}^+ + A^2 \cos^2 \theta} - A^2 \sin^2 \theta (r_{AH}^+ + A^2 \cos^2 \theta) \right] - \frac{1}{4} (r_{AH}^+ + A^2 \cos^2 \theta) \\ & \times \left[2 \arccos \theta + \frac{2m(r_{AH}^+ + A^2 \cos^2 \theta) - 4mr_{AH}^+}{(r_{AH}^+ + A^2 \cos^2 \theta)^2} + 2a^2 \sin^2 \theta r_{AH}^+ \right]. \end{aligned} \quad (52)$$

将(52)式代入(45)式整理后得

$$\begin{aligned} & -a^2 \sin^2 \theta r^4 - 2 \arccos \theta r^3 - (2A^2 a^2 \sin^2 \theta \cos^2 \theta + 2\dot{r} - 1) r^2 \\ & - (2m + 2A^2 \cos^3 \theta a) r - 2\dot{r} A^2 + A^2 - a^2 \sin^2 \theta A^4 \cos^4 \theta = 0. \end{aligned} \quad (53)$$

解(53)式得4个根,其中1个是负根无物理意义,其余3个根为

$$r_{EH_1} = -D'' + E'' + R'' - \frac{1}{2} \frac{\cos \theta}{a \sin^2 \theta}, \quad (54)$$

$$r_{EH_2} = D'' - E'' + R'' - \frac{1}{2} \frac{\cos \theta}{a \sin^2 \theta}, \quad (55)$$

$$r_{EH_3} = D'' + E'' - R'' - \frac{1}{2} \frac{\cos \theta}{a \sin^2 \theta}, \quad (56)$$

其中

$$D'' = \frac{1}{\sqrt{6} a \sin \theta} [1 - \alpha'' \cos(\varphi''/3 - \pi/3)]^{1/2}, \quad (57)$$

$$E'' = \frac{1}{\sqrt{6} a \sin \theta} [1 - \alpha'' \cos(\varphi''/3 + \pi/3)]^{1/2}, \quad (58)$$

$$R'' = \frac{1}{\sqrt{6} a \sin \theta} [1 + \alpha'' \cos(\varphi''/3)]^{1/2}, \quad (59)$$

$$\alpha'' = \left[\left(1 - r\dot{r} - 2A^2 a^2 \sin^2 \theta \cos^2 \theta + \frac{3}{2} \cot^2 \theta \right)^2 - 12 \left(\frac{3}{16} \frac{\cos^4 \theta}{a^2 \sin^6 \theta} + \frac{m \cos \theta}{a \sin^2 \theta} + \frac{1}{4} \frac{\cos^2 \theta}{a^2 \sin^4 \theta} + \frac{1}{2} \frac{A^2 \cos^4 \theta}{\sin^2 \theta} - 2\dot{r} A^2 + A^2 - a^2 \sin^2 \theta A^4 \cos^4 \theta \right) \right]^{1/2}, \quad (60)$$

$$\varphi'' = \arccos(-\beta''/\alpha'^3). \quad (61)$$

当 $0 < \varphi'' < \pi$ 时,

$$\begin{aligned} \beta'' = & \left(1 - 2\dot{r} - 2A^2 a^2 \sin^2 \theta \cos^2 \theta + \frac{3}{2} \cot^2 \theta \right)^3 + 30 a^2 \sin^2 \theta \left(1 - 2\dot{r} \right. \\ & \left. - 2A^2 a^2 \sin^2 \theta \cos^2 \theta + \frac{3}{2} \cot^2 \theta \right) \left(\frac{3}{16} \frac{\cos^4 \theta}{a^2 \sin^4 \theta} + \frac{m \cos \theta}{a \sin^2 \theta} + \frac{1}{2} \frac{A^2 \cos^4 \theta}{\sin^2 \theta} - 2\dot{r} A^2 \right. \\ & \left. + A^2 - A^4 a^2 \sin^2 \theta \cos^4 \theta \right) - 27 \left(2m + \frac{\cos \theta}{a \sin^2 \theta} + \frac{3}{2} \frac{\cos^3 \theta}{a \sin^4 \theta} \right)^2 a^2 \sin^2 \theta. \end{aligned} \quad (62)$$

r_{EH_1} 为 Rindler 视界, r_{EH_2} 和 r_{EH_3} 分别为匀加速直线运动的 Kerr 黑洞的外部 and 内部事件视界. 当 $m=0, A=0, \dot{r}=0$ 时, 得 Rindler 事件视界, 由方程(53)得

$$a^2 \sin^2 \theta r^2 + 2 a \cos \theta r - 1 = 0. \quad (63)$$

解之得

$$r = [a(1 + \cos \theta)]^{-1}. \quad (64)$$

恰为 Rindler 事件视界. 由此可知, 对于 Rindler 观测者而言, 它的 3 个视界面是重合的.

当 $a=0$ 时, 方程(53)变为

$$(1 - 2\dot{r}) r^2 - 2mr + (1 - 2\dot{r}) A^2 = 0, \quad (65)$$

解之得

$$r_{\text{EH}}^{\pm} = \frac{m \pm \sqrt{m^2 - A^2(1 - 2\dot{r})^2}}{(1 - 2\dot{r})}. \quad (66)$$

当 $\dot{r}=0$ 时, 事件视界回到熟知的稳态 Kerr 视界的位置,

$$r_{\text{EH}}^{\pm} = m \pm \sqrt{m^2 - A^2}. \quad (67)$$

总之, 我们求出了作匀加速直线运动的 Kerr 黑洞类时极限面、表观视界和事件视界的位置, 它们不再重合, 在动态情况下, 黑洞的外能层变厚, 内能层变薄.

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附 录

为了求表观视界, 我们计算出了不为零的联络为

$$\begin{aligned} \Gamma_{00}^0 &= -\frac{A^2 \sin^2 \theta}{2 p^2} \frac{dF}{dv} - \frac{1}{2} \left(1 + \frac{A^2}{p^2} \sin^2 \theta \right) \frac{dF}{dr}, \\ \Gamma_{02}^0 &= \Gamma_{20}^0 = - \left(1 + \frac{A^2}{p^2} \sin^2 \theta \right) r f - \frac{A^2 F}{2 p^2} \sin 2\theta, \end{aligned}$$

$$\begin{aligned}
\Gamma_{03}^0 &= \Gamma_{30}^0 = \frac{1}{2} \left(1 + \frac{A^2}{p^2} \sin^2 \theta \right) \left(A \sin^2 \theta \frac{dF}{dr} \right) + \frac{A^3 \sin^4 \theta}{2 p^2} \frac{dF}{dv}, \\
\Gamma_{12}^0 &= \Gamma_{21}^0 = -\frac{A^2 \sin 2\theta}{2 p^2}, \quad \Gamma_{13}^0 = \Gamma_{31}^0 = \frac{A r \sin^2 \theta}{p^2}, \\
\Gamma_{22}^0 &= -\frac{(A^2 + r^2) r}{p^2} - A^2 f \sin 2\theta, \\
\Gamma_{23}^0 &= \Gamma_{32}^0 = \frac{A^3 \sin^2 \theta}{2 p^2} F \sin 2\theta + \frac{(A^2 + r^2)}{p^2} A f r \sin^2 \theta, \\
\Gamma_{33}^0 &= -\frac{A^4 \sin^6 \theta}{2 p^2} \left(\frac{dF}{dv} + \frac{dF}{dr} \right) - \frac{1}{2} A^2 \sin^4 \theta \frac{dF}{dr} - \frac{(A^2 + r^2) r \sin^2 \theta}{p^2}, \\
\Gamma_{00}^1 &= \frac{1}{2} \left(1 - \frac{A^2}{p^2} \sin^2 \theta \right) + \frac{dF}{dv} + \frac{1}{2} g^{11} \frac{dF}{dr} + \frac{1}{2} f \frac{dF}{d\theta}, \\
\Gamma_{01}^1 &= \Gamma_{10}^1 = \frac{1}{2} \frac{dF}{dr} - r f^2, \\
\Gamma_{02}^1 &= \Gamma_{20}^1 = \frac{1}{2} \frac{dF}{d\theta} + g^{11} (rf) - \frac{A^2}{2 p^2} F \sin 2\theta, \\
\Gamma_{03}^1 &= \Gamma_{30}^1 = -\frac{1}{2} g^{11} \left(A \sin^2 \theta \frac{dF}{dr} \right) - \frac{1}{2} A f \sin^2 \theta \frac{dF}{d\theta} - \frac{1}{2} A f F \sin 2\theta + \frac{A^3 \sin^4 \theta}{2 p^2} \frac{dF}{dv}, \\
\Gamma_{12}^1 &= \Gamma_{21}^1 = -\frac{A^2 \sin 2\theta}{2 p^2}, \\
\Gamma_{13}^1 &= \Gamma_{31}^1 = -\frac{1}{2} A \sin^2 \theta \frac{dF}{dr} + A r f^2 \sin^2 \theta - A f \sin \theta \cos \theta + \frac{A r \sin^2 \theta}{p^2}, \\
\Gamma_{22}^1 &= f p^2 \cot \theta - \frac{3}{2} A^2 f \sin 2\theta + r g^{11}, \\
\Gamma_{23}^1 &= \Gamma_{32}^1 = -\frac{1}{2} A \sin^2 \theta \frac{dF}{d\theta} - \frac{1}{2} f^2 p^2 A \sin 2\theta - g^{11} A f r \sin^2 \theta, \\
\Gamma_{33}^1 &= -\frac{1}{2} A^2 \sin^4 \theta \left(1 + \frac{A^2}{p^2} \sin^2 \theta \right) \frac{dF}{dv} + \frac{1}{2} g^{11} \left(A^2 \sin^4 \theta \frac{dF}{dr} + 2 r \sin^2 \theta \right) \\
&\quad + \frac{1}{2} f \left(A^2 \sin^4 \theta \frac{dF}{d\theta} + 2 A^2 F \sin^2 \theta \sin 2\theta + (A^2 + r^2) \sin 2\theta \right), \\
\Gamma_{00}^2 &= \frac{1}{2} f \frac{dF}{dr} - \frac{1}{2 p^2} \frac{dF}{d\theta}, \\
\Gamma_{01}^2 &= \Gamma_{10}^2 = \frac{r f}{p^2}, \quad \Gamma_{02}^2 = \Gamma_{20}^2 = r f^2, \\
\Gamma_{03}^2 &= \Gamma_{30}^2 = -\frac{1}{2} f A \sin^2 \theta \frac{dF}{dr} + \frac{A \sin^2 \theta}{2 p^2} \frac{dF}{d\theta} + \frac{A F}{2 p^2} \sin 2\theta, \\
\Gamma_{12}^2 &= \Gamma_{21}^2 = \frac{r}{p^2}, \quad \Gamma_{13}^2 = \Gamma_{31}^2 = \frac{A \sin 2\theta - 2 A r f \sin^2 \theta}{2 p^2}, \\
\Gamma_{22}^2 &= r f - \frac{A^2 \sin 2\theta}{2 p^2}, \quad \Gamma_{23}^2 = \Gamma_{32}^2 = \frac{1}{2} A f \sin 2\theta - A f^2 r \sin^2 \theta, \\
\Gamma_{33}^2 &= \frac{1}{2} f \left(A^2 \sin^4 \theta \frac{dF}{dr} + 2 r \sin^2 \theta \right) - \frac{1}{2 p^2} \left(A^2 \sin^4 \theta \frac{dF}{d\theta} \right. \\
&\quad \left. + 2 A^2 F \sin^2 \theta \sin 2\theta + (A^2 + r^2) \sin 2\theta \right), \\
\Gamma_{00}^3 &= -\frac{A}{2 p^2} \left[\frac{dF}{dr} + \frac{dF}{dv} \right], \quad \Gamma_{02}^3 = \Gamma_{20}^3 = -\frac{A}{p^2} (rf + F \cot \theta), \\
\Gamma_{03}^3 &= \Gamma_{30}^3 = \frac{A^2 \sin^2 \theta}{2 p^2} \left(\frac{dF}{dr} + \frac{dF}{dv} \right), \quad \Gamma_{12}^3 = \Gamma_{21}^3 = -\frac{A}{p^2} \cot \theta, \\
\Gamma_{13}^3 &= \Gamma_{31}^3 = \frac{r}{p^2}, \quad \Gamma_{22}^3 = -2 A f \cot \theta - \frac{A r}{p^2},
\end{aligned}$$

$$\Gamma_{23}^3 = \Gamma_{32}^3 = \frac{(A^2 + r^2)}{p^2} \cot \theta + \frac{A^2 f r \sin^2 \theta}{p^2} + \frac{A^2 F \sin 2\theta}{2 p^2},$$

$$\Gamma_{33}^3 = -\frac{A^3 \sin^4 \theta}{2 p^2} \left(\frac{dF}{dv} + \frac{dF}{dr} \right) - \frac{Ar}{p^2} \sin^2 \theta.$$

- [1] W. A. Hiscock, *Phys. Rev.*, **D23**(1981), 2813.
- [2] R. Balbinot A. Barletta, *Class. Quantum Grav.*, **6**(1989), 195.
- [3] R. Balbinot, *Phys. Rev.*, **D33**(1986), 1611.
- [4] 宋世学、刘 辽、赵 峰, 北京师范大学学报(自然科学版), **29**(3)(1993), 357.
- [5] R. L. Mallett, *Phys. Rev.*, **D33**(1986), 2021.
- [6] Yu Hongwei, *Phys. Lett.*, **A209**(1995), 6.
- [7] Jing Ji-liang, Yu Hong-wei, Wang Yong-jiu, *Chin. Phys. Lett.*, **9**(1992), 113.
- [8] J. W. York, Jr., *Quantum Theory of Gravity*, ed. by S. Christensen (1984), p. 135.

CHARACTERISTIC SURFACES OF A UNIFORMLY RECTILINEARLY ACCELERATING KERR BLACK HOLE

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ABSTRACT

In this paper, we deal with three kinds of characteristic surfaces of a uniformly rectilinearly accelerating Kerr black hole whose mass is varying. i. e., the event horizon, the apparent horizon, and the time-like limit surface (the infinite red-shift surface). It is pointed out that the three kinds of characteristic surfaces no longer coincide with one other when the black hole mass is varying (accreting or evaporating).

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